

Kalman filtering is super cool!

Scalar case setup:

$$x_n = ax_{n-1} + v_n, \quad v_n \sim \mathcal{N}(0, \sigma_v^2)$$

$$y_n = bx_n + w_n, \quad w_n \sim \mathcal{N}(0, \sigma_w^2)$$

states

measurements

Want  $\hat{x}_n \triangleq E[x_n | y_1, y_2, \dots, y_n]$

Intuition: two noisy estimates of  $x_n$   
Some noise is "teeth", some is just from the measurement.

① The physics estimate:

given my MMSE state so far ( $\hat{x}_{n-1}$ )  
what's going to happen next?

$$x_{n, \text{phy}} = E[x_n | \hat{x}_{n-1}] = a \hat{x}_{n-1}$$

② The measurement estimate:

what state corresponds most closely with the  
observation  $y_n$ ?

$$x_{n, \text{mes}} = \frac{1}{b} y_n$$

→ this doesn't exactly generalize to  
the vector case: with  $y_n = Bx_n + w_n$ ,  
 $B$  is often not square/invertible.

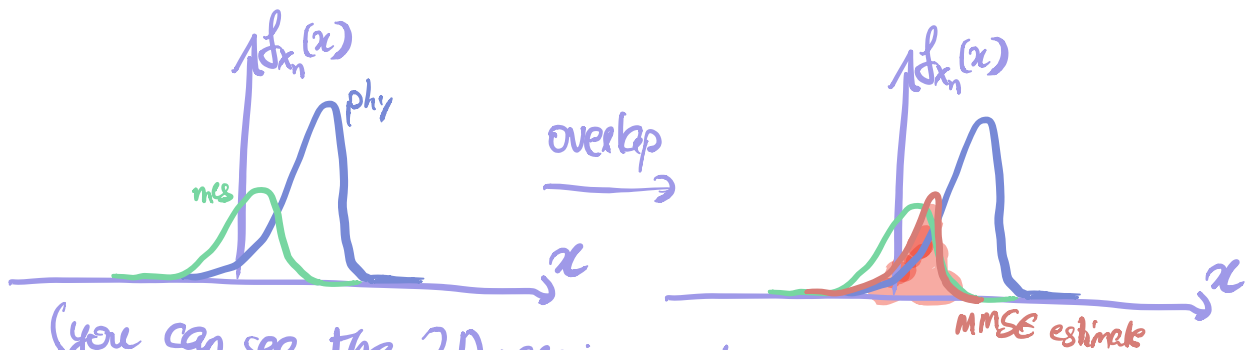
Want to make an overall MMSE estimate.

Intuitively, this is some linear combination of

$X_{n,phy}$  and  $X_{n,mes}$ .

$$\hat{X}_n = \gamma X_{n,phy} + (1-\gamma) X_{n,mes}$$

Find  $\gamma$  and we're done!



(you can see the 2D version of this in the KF lab!)

How do we do this overlap in algebra?

Since all variables are  $IG$ , LLSE = MMSE,  
so let's frame this as an LLSE update question!

(upcoming new notation: LLSE update only holds  
for zero-mean, so we'll say  $\bar{X}_n = X_n - E[X_n]$ )

Some dense algebra...

$$\hat{x}_n = E[x_n | y_1, y_2, \dots, y_n]$$

$$= L[x_n | y_1, y_2, \dots, y_n] \quad (\text{JG})$$

$$= L[x_n | y_1] + L[\bar{x}_n | y_2 - L[y_2 | y_1]] + L[\bar{x}_n | y_3 - L[y_3 | y_1, y_2]] \\ + \dots + L[\bar{x}_n | y_n - L[y_n | y_1, y_2, \dots, y_{n-1}]]$$

(USE update)

$$= a L[x_{n-1} | y_1] + a L[\bar{x}_{n-1} | y_2 - L[y_2 | y_1]] + a L[\bar{x}_{n-1} | y_3 - L[y_3 | y_1, y_2]] \\ + \dots + L[\bar{x}_n | y_n - L[y_n | y_1, y_2, \dots, y_{n-1}]]$$

(if you don't see  $y_n$ , there's no update info, so  
your best guess for  $x_n = a \cdot$  your best guess for  $x_{n-1}$ )

$$= a L[x_{n-1} | y_1, y_2, \dots, y_{n-1}] + L[\bar{x}_n | y_n - L[y_n | y_1, y_2, \dots, y_{n-1}]]$$

(collected all the  $a L[x_{n-1} | \dots]$  terms)

$$= a \hat{x}_{n-1} + L[\bar{x}_n | y_n - b L[x_n | y_1, y_2, \dots, y_{n-1}]]$$

$$= a \hat{x}_{n-1} + L[\bar{x}_n | y_n - ab L[x_{n-1} | y_1, y_2, \dots, y_{n-1}]]$$

$$= a \hat{x}_{n-1} + k_n (y_n - ab \hat{x}_{n-1})$$

this step from

"Kalman gain"      "innovation" } → LLSE being linear  
 → update is zero-mean  
 $\tilde{Y}_n \cong Y_n - ab\hat{X}_{n-1}$

Note here:

$$\begin{aligned} \mathcal{L}[X_n | Y^{(n)}] &= \mathcal{L}[X_n | Y^{(n-1)}] \\ &\quad + \mathcal{L}[X_n | \tilde{Y}_n] \end{aligned}$$

→  $k_n = \frac{\text{cov}(X_n, \tilde{Y}_n)}{\text{var} \tilde{Y}_n}$  by LLSE formula

→ innovation  $\tilde{Y}_n$  is zero-mean:  
 in expectation we have just  $X_n = a^n X_0$ .

How do we find  $k_n$ ? Slightly complicated, because it's from a sequence of Gaussian overlaps:

denote a Gaussian overlap by  $\otimes$

$$\begin{aligned} \hat{X}_1 &\sim \mathcal{N}(X_0, \sigma_v^2) \otimes \mathcal{N}\left(\frac{1}{b} Y_1, \frac{\sigma_w^2}{b^2}\right) \\ \hat{X}_2 &\sim \mathcal{N}(\hat{X}_1, \sigma_v^2) \otimes \mathcal{N}\left(\frac{1}{b} Y_2, \frac{\sigma_w^2}{b^2}\right) \\ &= \mathcal{N}\left(\mathcal{N}(X_0, \sigma_v^2) \otimes \mathcal{N}\left(\frac{1}{b} Y_1, \frac{\sigma_w^2}{b^2}\right), \sigma_v^2\right) \otimes \mathcal{N}\left(\frac{1}{b} Y_2, \frac{\sigma_w^2}{b^2}\right) \\ X_3 &\sim \dots \end{aligned}$$

This is a lot!

this was a discussion question at some point

Key here: overlap of Gaussians is Gaussian, so let's keep track of the sequence of estimator variances.

Say  $\hat{X}_n \sim \mathcal{N}(X_n, \sigma_{n|n}^2)$ .

(intuitively, e.g.  $P(|\hat{X}_n - X_n| < 2\sigma_{n|n}) > 0.95$ .)

and say  $\sigma_{n|n-1}^2 = a^2 \sigma_{n-1|n-1}^2$  - variance just of the physics estimate.

$$k_n = \frac{\text{cov}(X_n, \tilde{Y}_n)}{\text{var}(\tilde{Y}_n)} = \frac{\text{cov}(X_n, X_n - ab\hat{X}_{n-1})}{\text{var}(\tilde{Y}_n)}$$

(I'm lazy, see website notes)

$$k_n = \frac{\sigma_{n|n-1}^2}{\sigma_{n|n-1}^2 + \sigma_w^2} = \frac{a^2 \sigma_{n-1|n-1}^2}{a^2 \sigma_{n-1|n-1}^2 + \sigma_w^2}$$

Lots of algebra for basically a weighted average on variances: rewrite the estimator in the form from the start -

$$\hat{X}_n = \gamma X_{n,\text{phy}} + (1-\gamma) X_{n,\text{mes}}$$

$$\hat{X}_n = a\hat{X}_{n-1} - abk_n\hat{X}_{n-1} + k_n Y_n$$

$$= (a - abk_n) X_{n,\text{phy}} + k_n b X_{n,\text{mes}}$$

This is one of the most practically applicable concepts in the course!

Example: I was in a rocketry club ([stars.berkeley.edu](mailto:stars.berkeley.edu))

We wanted to know how high the rocket went over time, because

→ knowing how high your rocket went is fun

→ it helps you time parachute release on the way down

Rocket state:  $\begin{bmatrix} z \\ v_z \\ a_z \end{bmatrix}$  and sensors give you  $N\left(\begin{bmatrix} z \\ a_z \end{bmatrix}, \Sigma_v\right)$ .

→ Sensors are an imperfect estimate

→ We also have physics estimates (F=ma: the model comes w/ a model of F(t), so that aren't perfect due to environmental noise (wind, etc) we get an ideal alt.)

MMSE rocket states come from Kalman filtering w/ the physics model + sensor measurements!

(You'll work out the exact algebra for a similar case to this in the lab!)

Key to the KF equations: recursion.

$$E[Y_n | Y_1, \dots, Y_{n-1}] = \alpha \beta \hat{X}_{n-1}$$

$$\begin{aligned} E[X_n | Y_1, \dots, Y_{n-1}] &= E[\alpha X_{n-1} + V_n | Y^{(n-1)}] \\ &= \alpha \hat{X}_{n-1} \end{aligned}$$